AN NONEXISTENCE THEOREM FOR HARMONIC MAPS WITH SLOWLY DIVERGENT ENERGY

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Abstract

It is proved that a harmonic map or a relative harmonic map from Euclidean space $R^n(n \neq 2)$ into an m-dimensional Riemannian manifold M_m with finite energy or slowly divergent energy must be a constant map. Some physical applications are also presented.

It is known that the harmonic map from Euclidean space $R^n(n>2)$ to any m dimensional Riemannian space M_m with finite energy must be a constant map, i. e. the image of R^n is a fixed point $^{(1-3)}$. In this paper, we will use a certain technique, similar to that in the author's previous work on massive Yang-Mills fields $^{(4)}$, to show that the "finite energy" hypothesis can be weakened. The method is quite different from that in the papers [1-3].

We also show that this conclusion for the harmonic maps holds for relative harmonic maps, too.

As is known, the energy of a harmonic map ϕ from $R^n \rightarrow M_m$ is

$$E(\phi) = \int_{\mathbb{R}^n} e(\phi) d^n x. \tag{1}$$

In local coordinates, the energy density $e(\phi)$ can be expressed in the form

$$e(\phi) = \sum_{i} g_{\alpha\beta}(\phi) \frac{\partial \phi^{\alpha}}{\partial x^{i}} \frac{\partial \phi^{\beta}}{\partial x^{i}} = \sum_{i} g_{\alpha\beta} \phi_{i}^{\alpha} \phi_{i}^{\beta}$$

$$(\alpha, \beta = 1, \dots, m; i, j = 1, \dots, n),$$

$$(2)$$

where $g_{\alpha\beta}$ is the metric tensor of M_m . The equations for harmonic maps are

$$\sum_{i} \phi_{i|i}^{\alpha} = 0. \tag{3}$$

where "|" denotes the covariant derivatives, i. e.

$$\phi_{i|i}^{\alpha} = \frac{\partial \phi_{i}^{\alpha}}{\partial x^{i}} + \Gamma_{\beta\gamma}^{\alpha} \phi_{i}^{\beta} \phi_{i}^{\gamma}. \tag{4}$$

Here $\Gamma^{\alpha}_{\beta\gamma}$ are the Christoffel symbols of M_m .

For each harmonic map ϕ , we have a symmetric tensor

$$T_{ij} = \frac{1}{2} \left(\sum_{i} g_{\alpha\beta} \phi_{i}^{\alpha} \phi_{i}^{\beta} \right) \delta_{ij} - g_{\alpha\beta} \phi_{i}^{\alpha} \phi_{j}^{\beta}, \tag{5}$$

which satisfies the conservative law^[5]

$$\frac{\partial T_{ij}}{\partial x^j} = 0. ag{6}$$

It is noted that (6) also follows from the Noether's theorem in the field theory.

In the previous works, it was assumed that the energy $E(\phi)$ of the harmonic map ϕ from $R^n \rightarrow M_m$ is finite, now the condition is weakened to

$$\int \frac{e(\phi)}{\psi(r)} d^n x < \infty, \tag{7}$$

where $\psi(r)$ is a positive, continuous function of r satisfying

$$\int_{a}^{\infty} \frac{dr}{r\psi(r)} = \infty \quad \text{(for a certain constant } a > 0\text{).}$$
 (8)

If $\psi(r) = 1$, the energy is finite. But the energy may be infinite, for example, in the case $\psi(r) = O(\log r)$ (as $r \to \infty$). Hence when (7) holds true, the energy $E(\phi)$ of a harmonic map ϕ may be either finite or infinite.

Difinition. The energy is called "slowly divergent" if $\int e(\phi) d^n x = \infty$ and (7) holds.

Theorem 1. Let $\phi: R^n \to M_m$ be a harmonic map from n(n>2) dimensional Euclidean space R^n into an m-dimensional Riemannian manifold M_m . Suppose that the energy $E(\phi)$ of ϕ is finite or slowly divergent, then ϕ is a constant map.

Proof For any piecewise continuous function $\sigma(r)$, we have

$$0 = \int_{0}^{R} \sigma(r) dr \int_{|x| < r} x^{j} \frac{\partial T_{ij}}{\partial x^{i}} d^{n}x$$

$$= \int_{0}^{R} \sigma(r) dr \int_{|x| < r} \left\{ \frac{\partial}{\partial x^{i}} (x^{j} T_{ij}) - \sum_{i} T_{ii} \right\} d^{n}x$$

$$= \int_{0}^{R} \sigma(r) dr \int_{|x| = r} (x^{j} T_{ij}) \frac{x^{i}}{r} dS - \int_{0}^{R} \sigma(r) dr \int_{|x| < r} \sum_{i} T_{ii} d^{n}x.$$

$$(9)$$

From (5), we obtain

$$\sum_{i} T_{ii} = \left(\frac{n}{2} - 1\right) \left(\sum_{i} g_{\alpha\beta} \phi_{i}^{\alpha} \phi_{i}^{\beta}\right) \geqslant 0.$$
 (10)

Suppose ϕ is not a constant harmonic map, then there exist positive constants R_1 and ϵ such that

$$\int_{|x| \leq R} \sum_{i} T_{ii} d^{n} x > \epsilon > 0 \tag{11}$$

for $R \geqslant R_1$.

Choose $\sigma(r) \geqslant 0$ such that $\sigma(r) = 0$ iff $r \leqslant R_1$. Then, from (9) we have

$$0 < \int_{0}^{R} \sigma(r) dr \int_{|x|=r} x^{j} T_{ij} \frac{x^{i}}{r} dS - \epsilon \int_{R_{i}}^{R} \sigma(r) dr.$$
 (12)

Since $|T_{ij}| \leq \frac{3}{2} e(\phi)$, we have

$$0 < \frac{3}{2} n^2 \int_{|x| \leq R} r \sigma(r) e(\phi) d^n x - \epsilon \int_{R_1}^R \sigma(r) dr.$$
 (13)

Furthermore, we choose $\sigma(r)$ such that for $r > R_1$

$$\sigma(r) = \frac{1}{r\psi(r)},\tag{14}$$

then from (7) and (8), it is easily seen that if R is sufficiently large, the right side of (13) should be negative. This is a contradiction. Hence ϕ must be a constant map. Theorem 1 is proved.

A relative harmonic map ϕ is a generalization of the harmonic map. It is defined in [6] by the formula

$$\sum_{k} g_{\alpha\beta} \phi_{k|k}^{\alpha} \phi_{k}^{\beta} = 0 \tag{15}$$

in local coordinates. We can verify that for relative harmonic maps, equation (6) holds true also. Repeating the previous argument, we have

Theorem 2. There exists no relative harmonic map from $R^n \rightarrow M_m(n>2)$ with finite energy or slowly divergent energy other than a constant map.

In theoretical physics, the Chiral field (or the nonlinear σ -model) in 4-dimensional Minkowski spacetime $R^{3,1}$ is just a harmonic map ϕ from $R^{3,1}$ to a homogeneous Riemannian manifold M_m . If the field is static, then ϕ is a harmonic map from R^3 to M_m . Hence the physical significance of Theorem 1 can be expressed as

Theorem 3. In n+1 (n>2) dimensional Minkowski spacetime $R^{n,1}$, there does not exist any static nontrivial Chiral field with finite energy or slowly divergent energy.

It is known that Ernst equations are the basic equations for obtaining axial symmetric stationary solutions to the pure Einstein gravitational equation and are the Euler equations of the energy integral

$$\int \frac{\sum_i \left[\left(\frac{\partial \varPhi}{\partial x^i} \right)^2 + \left(\frac{\partial \varPsi}{\partial x^i} \right)^2 \right]}{\varPhi^2} d^n x.$$

Hence a solution to the Ernst equations is a harmonic map from \mathbb{R}^n to the hyperbolic plane with metric

$$dS^2 = \frac{1}{\Phi^2} (d\Phi^2 + d\Psi^2), \quad \Phi > 0$$

in Poincare representation.

Thus we have

Theorem 4. A nontrivial solution to the Ernst equations with n>2 must have infinite energy. Furthermore, the energy cannot be slowly divergent. Here the energy is in the sense of harmonic maps.

Remark 1. There are many nontrivial harmonic maps ϕ from S^2 to certain manifolds with finite energy $E(\phi)$. By using sterographic projection from $S^2 \rightarrow R^2$, we will obtain nontrivial harmonic maps from $R^2 \rightarrow M$ with finite energy. So n=2 is actually an exceptional case.

Remark 2. We shall specialize that the condition for the energy in our theorem cannot be omitted, because we can find many regular harmonic maps from R^n , whose energy does not diverge so slowly. For example, the linear function

 $\phi = a_i x^i + a$ (a_i are constants not all zero)

is a harmonic map from R^n to R with a constant energy density.

Further results on harmonic maps from noncompact Riemannian manifolds will appear in [7].

References

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